

1. Introduction

• Heavy Ion Collisions and Quark-Gluon Plasma

- **Quark-gluon plasma (QGP)**, a deconfined state of quarks and gluons, is produced in **heavy ion collision** at RHIC and LHC.
- The QGP in these experiments behaves like a **strongly coupled liquid**, not like a weakly coupled gas.
- The plasma **thermalizes on a very short time scale** of $\approx 10^{-23}$ sec, which is theoretically not well understood yet.
- Due to strong coupling **perturbative QCD is not suitable** to study the **quantum dynamics** of these collisions.

• AdS/CFT Correspondence [1]

- AdS/CFT maps strongly coupled supersymmetric Yang-Mills (SYM) theory in 4D to classical gravity on 5D anti-de Sitter (AdS5) space.
- We use SYM theory as a **toy model** for QCD.
- **Thermalization** in the strongly coupled 4D SYM theory is mapped to **black hole formation** in AdS5.

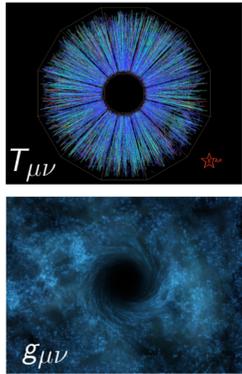


Figure 1: AdS/CFT maps thermalization in the 4D gauge theory (top) to black hole formation in a 5D gravity theory (bottom).

3. Shock Wave Collisions in SYM theory

The **Lorentz contracted "nuclei"** in SYM are modelled as **two Gaussian energy distributions** approaching each other at the **speed of light**. The **time evolution** of the energy-momentum tensor is extracted from a **numerical relativity simulation** of **colliding gravitational shock waves** in the 5D gravity theory. [4]

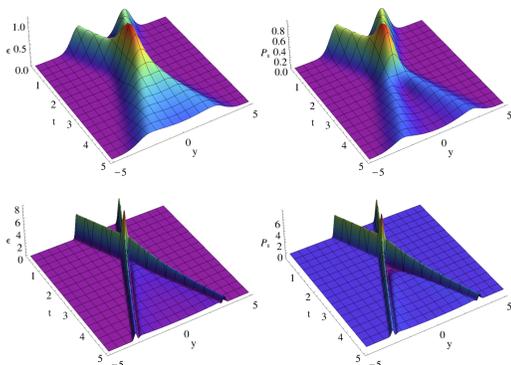


Figure 3: Energy density \mathcal{E} (left) and longitudinal pressure \mathcal{P}_{\parallel} (right) for wide (top) and narrow shocks (bottom).

• Energy-Momentum Tensor

$$\langle T^{\mu\nu} \rangle = \frac{N^2}{2\pi^2} \begin{pmatrix} \mathcal{E} & \mathcal{S} & 0 & 0 \\ \mathcal{S} & \mathcal{P}_{\parallel} & 0 & 0 \\ 0 & 0 & \mathcal{P}_{\perp} & 0 \\ 0 & 0 & 0 & \mathcal{P}_{\perp} \end{pmatrix}$$

- **Wide and narrow shocks** show **qualitatively different behavior**. [5]

• Wide Shocks: Full-Stopping

- Wide shocks **stop each other** at the collision when the plasma is formed which then **explodes hydrodynamically**.
- **Energy and pressure stay positive**.

• Narrow Shocks: Transparency

- Narrow shocks **pass each other** almost **"transparently"** and the plasma is formed only after the collision.
- **Energy and pressure** can be **negative** for a short time period after the collision.

5. Two-Point Functions

Time Evolution Two-Point Functions

- The system **starts** in some **correlated state**.
- As the **shocks approach** each other **without interaction** they **destroy** these initial correlations.
- **After the collision** correlations are **restored** because of the **interactions during the collisions** new correlations are formed.
- For the **narrow shocks** these new correlations **grow significantly beyond their initial value**.
- The shock wave system follows a **top-down thermalization** pattern where **short range correlations** (small L) **reach the equilibrium first**.

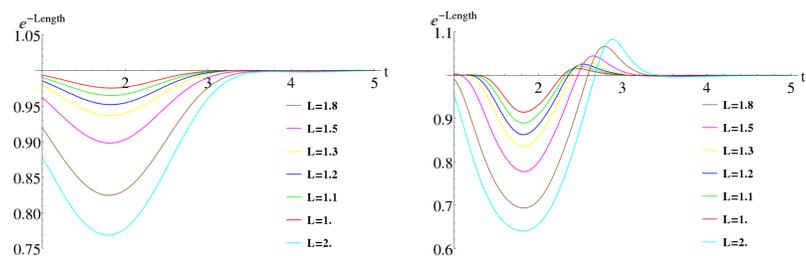


Figure 5: Two-point function of various separations L for wide shocks (left) and narrow shocks (right).

7. Summary

- We use collisions of **shock waves in SYM theory** as **toy model for real (QCD) heavy ion collisions**.
- Using the **AdS/CFT correspondence** the dynamics in these collisions can be extracted from **numerical relativity simulations of colliding gravitational shock waves**.
- Within AdS/CFT **non-local observables** such as **two-point functions** and **entanglement entropy** can be computed from **geodesics** and **minimal surfaces** in the gravity theory.
- From our numerical simulation we find that both, **two-point functions and entanglement entropy**, show **qualitatively different behavior for narrow and wide shocks**.
- A Mathematica code for shock wave collisions is available at Wilke van der Schee's homepage: www.sites.google.com/site/wilkevanderschee/phd-thesis
- A Mathematica code for the entanglement entropy and the two-point function can be downloaded from: www.christianecker.com

2. Local and Non-Local Observables from AdS/CFT

Within AdS/CFT it is possible to compute **expectation values of observables** in 4D SYM theory from **purely geometric objects** in the 5D gravity theory such as the **metric**, **geodesics** and **minimal surfaces**.

• Local Observables

- The **energy-momentum tensor** can be extracted from the **metric** $g_{\mu\nu}$ near the boundary.

$$\langle T^{\mu\nu}(x) \rangle = -\frac{2}{\text{det}g} \frac{\delta S}{\delta g^{\mu\nu}(x)}$$

• Non-Local Observables

- **Two-point functions** for gauge invariant operators \mathcal{O} of large conformal weight Δ are given by the **length of geodesics** γ . [2]

$$\langle \mathcal{O}(t, \vec{x}) \mathcal{O}(t, \vec{x}') \rangle \approx e^{-\Delta \text{Length}(\gamma)}$$

- **Entanglement entropy** of a spatial region A is given by the **area of a minimal surface** Σ . [3]

$$S_A = -\text{Tr}_{\rho_A} \log \rho_A = \frac{\text{Area}(\Sigma)}{4G_N}$$

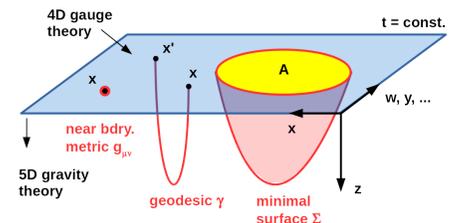


Figure 2: Geometric description of the energy momentum tensor, two-point functions and entanglement entropy in terms of the near boundary metric, geodesics and minimal surfaces.

4. Geodesics in AdS5

In the calculation of **two-point functions** we need the **length of spacelike geodesics** that are **attached to the boundary** at $z=0$ and extend into the 5D shock wave geometry.

These geodesics can be found by **numerically solving the geodesic equation** (1) subject to **boundary conditions** (2) that fix the endpoints at the boundary at some **spatial separation** L . [6]

$$\ddot{X}^\mu + \Gamma^\mu_{\alpha\beta} \dot{X}^\alpha \dot{X}^\beta = -J \dot{X}^\mu \quad (1)$$

$$X^\mu(\sigma_\pm) \equiv (V(\sigma_\pm), Z(\sigma_\pm), Y(\sigma_\pm)) = (t, z_{cut}, \pm L/2) \quad (2)$$

During the collision a **black hole horizon** is formed in the 5D shock wave geometry. The horizon has **different shape** for wide and narrow shock waves. The **geodesics tend not to cross the horizon** which leads to distortions in their **shape and length** that is **characteristic** for wide and narrow shocks.

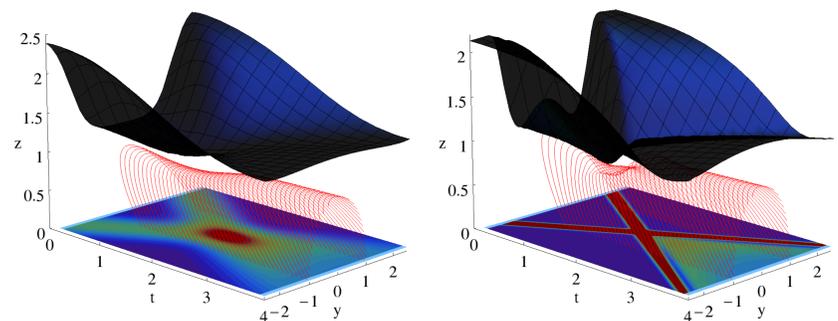


Figure 4: Black hole horizon (black), geodesics (red) and energy contours at $z=0$ for wide (left) and narrow shocks (right).

6. Entanglement Entropy

Time Evolution of Entanglement Entropy

- We **start with zero entanglement** by construction.
- As the **shocks enter the entangling region** the **entanglement entropy rapidly grows**.
- After the rapid initial growth follows a regime of **linear growth** which goes approx. **until the shocks collide**.
- **Right after the collision** the **wide shocks** show a **smooth fall** of where the **narrow shocks** have a **pronounced minimum** which is related to the **minima in the energy density** and the **longitudinal pressure**.
- As the two-point function the **entanglement entropy** shows a **top-down thermalization** pattern.

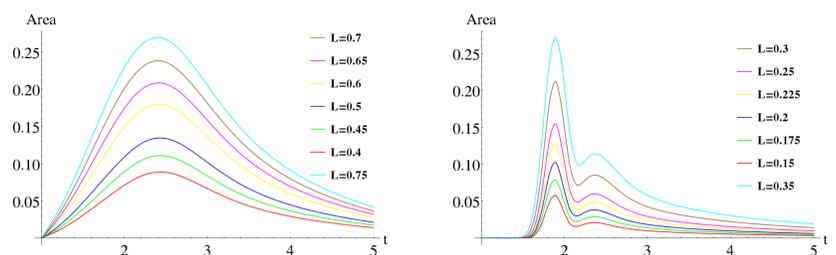


Figure 6: Entanglement entropy of various system sizes L for wide shocks (left) and narrow shocks (right).

8. References

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